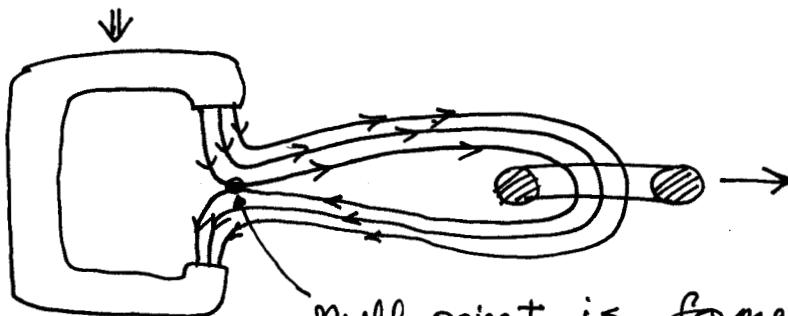
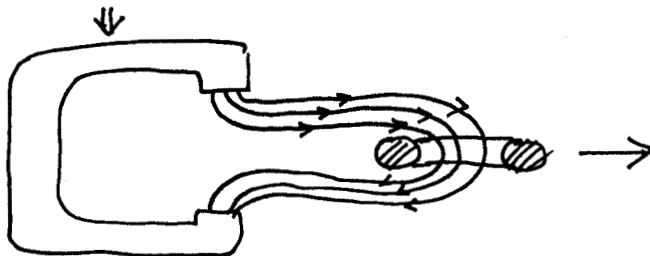
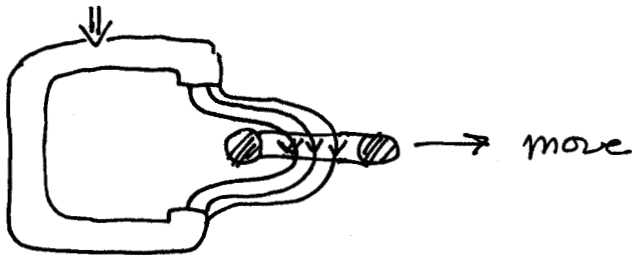
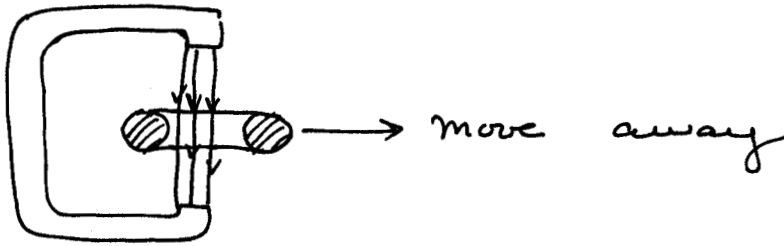
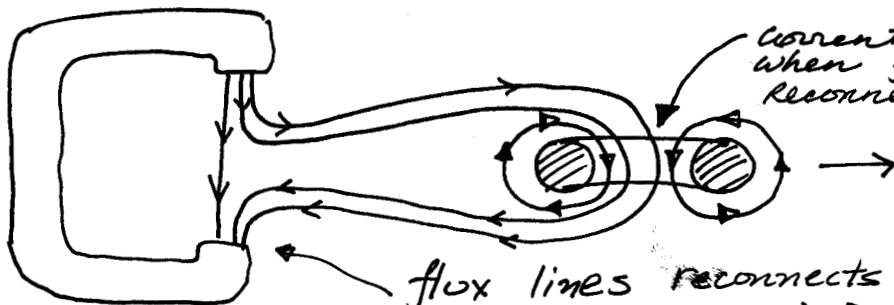


1)

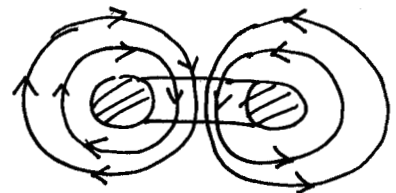
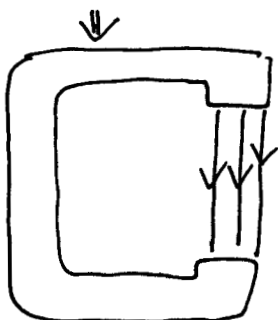


Null point is formed! (i.e. $B=0$)



Current is induced when flux lines reconnects!

flux lines reconnects at null point, hence $\nabla \cdot \vec{B}$ always satisfied.



Notes:

→ Reconnection occurs only w/ resistivity (you'll learn about that later!)

→ the flux through the torus is constant thru the whole process! (Current is induced as required.)

2. Polarization Drift

We now've $\vec{v}_{\text{EXB}} = \vec{v}_{\text{EXB}}(t)$

Using conservation of energy,

$$\frac{d}{dt} \left(\frac{1}{2} m \vec{v}_{\text{EXB}}^2 \right) = \vec{v}_{\text{pol}} \cdot q \vec{E}_{\perp}(t)$$

$$\vec{v}_{\text{EXB}} = \frac{\vec{E}_{\perp} \times \vec{B}}{B^2}$$

$$v_{\text{EXB}}^2 = \frac{E_{\perp}^2}{B^2}$$

then,

$$\frac{d}{dt} \left(\frac{1}{2} m \frac{E_{\perp}^2}{B^2} \right) = \frac{1}{2} \frac{m}{B^2} \cancel{2 \vec{E}_{\perp}} \cdot \frac{d\vec{E}_{\perp}}{dt} = \vec{v}_{\text{pol}} \cdot q \vec{E}_{\perp}(t)$$

then

$$\boxed{\vec{v}_{\text{pol}} = \frac{m}{qB^2} \frac{d\vec{E}_{\perp}}{dt} = \frac{1}{\Omega B} \frac{d\vec{E}_{\perp}}{dt}}$$

→ that's the quick way

→ the rigorous way uses
gyro-average theory →

2) Polarization Drift (Cont)

• Let's start w/ basic EOH

$$m \frac{d\vec{v}}{dt} = q(\vec{E}(t) + \vec{v} \times \vec{B})$$

• choose $\vec{B} = B_0 \hat{z}$ & $\vec{E} = E(t) \hat{x}$

then we've

$$(1) \quad m \frac{dv_x}{dt} = q E_x(t) + q v_y B_0$$

$$(2) \quad m \frac{dv_y}{dt} = -q v_x B_0$$

taking a derivative of (1) & (2) & substituting,
we've

$$m \frac{d^2 v_x}{dt^2} = q \frac{\partial E_x}{\partial t} + q \left(-\frac{q}{m} v_x B_0 \right) B_0$$

&

$$m \frac{d^2 v_y}{dt^2} = -q \left(\left(\frac{q}{m} \right) (E_x(t) + v_y B_0) \right) B_0$$

Rewriting,

$$(3) \quad \frac{d^2 v_x}{dt^2} = \frac{\Omega}{B_0} \frac{\partial E_x}{\partial t} - \Omega^2 v_x$$

$$\frac{d^2 v_y}{dt^2} = -\Omega \left(\frac{\Omega}{B_0} E_x + \Omega v_y \right)$$

$$(4) \quad \frac{d^2 v_y}{dt^2} = -\frac{\Omega^2}{B_0} E_x - \Omega^2 v_y$$

2) Cont

Consider $\vec{v} = \vec{v}_0 + \vec{v}_1 + \vec{v}_2$,

where v_0 is the gyro-motion

v_1 the standard $E \times B$ drift

v_2 the polarization drift

then,

$$\frac{d^2(\vec{v}_0 + \vec{v}_1 + \vec{v}_2)}{dt^2} = \underbrace{\frac{\Omega}{B_0} \frac{\partial \vec{E}_x}{\partial t}}_{2^{nd} \text{ order}} - \underbrace{\frac{\Omega^2}{B_0} E_x \hat{y}}_{1^{st} \text{ order}} - \Omega^2 (\vec{v}_0 + \vec{v}_1 + \vec{v}_2)$$

so, we've got

• 0th order $\frac{d^2 \vec{v}_0}{dt^2} = -\Omega^2 \vec{v}_0$, which is zeroth order gyro-motion

(if u take the gyro-average $\frac{d^2 \langle \vec{v}_0 \rangle}{dt^2} = 0$ so $\langle \vec{v}_0 \rangle = 0!$)
 ↑ Guiding center

• 1st order $\frac{d^2 v_{1y}}{dt^2} = -\frac{\Omega^2}{B_0} E_x - \Omega^2 v_{1y}$

taking the gyro-average, we've

$$\frac{d^2 \langle v_{1y} \rangle}{dt^2} = -\frac{\Omega^2}{B_0} \langle E_x \rangle - \Omega^2 \langle v_{1y} \rangle$$

then, $\frac{d^2 \langle v_{1y} \rangle}{dt^2} \ll -\Omega^2 \langle v_{1y} \rangle$

so

$$v_{1y} = -\frac{\langle E_x \rangle}{B_0} \Rightarrow \text{which is just the } \frac{E \times B}{B^2} \text{ drift}$$

• 2nd order $\frac{d^2 \langle v_{2x} \rangle}{dt^2} = \frac{\Omega}{B_0} \frac{\partial \langle E_x \rangle}{\partial t} - \Omega^2 \langle v_{2x} \rangle$

again $\frac{d^2 \langle v_{2x} \rangle}{dt^2} \ll -\Omega^2 \langle v_{2x} \rangle$,

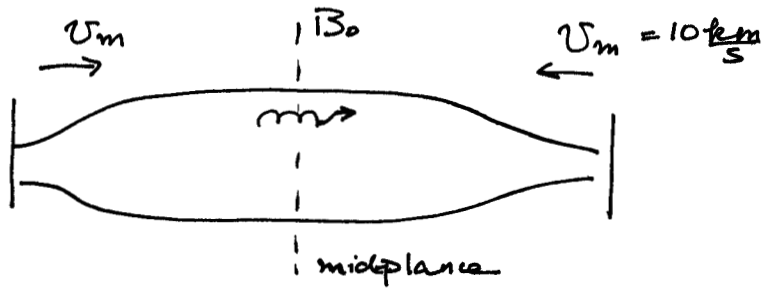
so

$$+\Omega^2 \langle v_{2x} \rangle = \frac{\Omega}{B_0} \frac{\partial \langle E_x \rangle}{\partial t}$$

$$\boxed{\langle v_{2x} \rangle = \frac{1}{\Omega B_0} \frac{\partial \langle E_x \rangle}{\partial t}}$$

the polarization drift!
 In general $\vec{v}_2 = \frac{1}{\Omega B_0} \frac{\partial \vec{E}}{\partial t}$

3) Fermi Acceleration of Cosmic Rays



• Let i & f represent the initial and final states (before & after acceleration)

a) • Since μ is constant both in time and space, we first see: for velocities at the midplane:

$$v_{\perp i} = v_{\perp f} \text{ at midplane, since } B_{0i} = B_{0f} \\ (\mu_i = \mu_f)$$

• Now, let's examine the situation of the μ constant in space at t_{final} .

So we've for the loss cone condition:

$$\frac{B_{0f}}{B_{mf}} = \sin^2 \theta_m = \frac{1}{R_m} = \frac{1}{5} = \frac{v_{\perp of}^2}{v_{\perp mf}^2}$$

• Using conservation of energy at t_{final} :

$$v_{\perp of}^2 + v_{\parallel of}^2 = v_{\perp mf}^2$$

then,

$$\frac{1}{5} = \frac{v_{\perp of}^2}{v_{\perp of}^2 + v_{\parallel of}^2}$$

• Using $v_{\perp of}^2 = v_{\perp oi}^2$

$$\frac{1}{5} = \frac{v_{\perp oi}^2}{v_{\perp oi}^2 + v_{\parallel of}^2} = \frac{1}{1 + \frac{v_{\parallel of}^2}{v_{\perp oi}^2}}$$

hence, $\frac{v_{\parallel of}^2}{v_{\perp oi}^2} = 4 \Rightarrow v_{\parallel of} = 2 v_{\perp oi}$

So Now let's look at the Energies at the mid-plane

$$E_i = \frac{1}{2} m (v_{\perp 0i}^2 + v_{\parallel 0i}^2) = \frac{2}{2} m (v_{\perp 0i}^2) = m v_{\perp 0i}^2$$

$$E_f = \frac{1}{2} m (v_{\perp 0f}^2 + v_{\parallel 0f}^2) = \frac{1}{2} m (v_{\perp 0i}^2 + 4 v_{\perp 0i}^2)$$

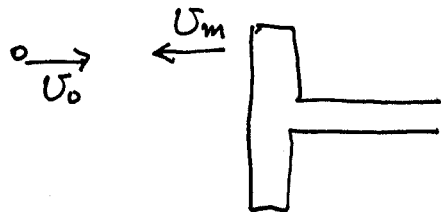
so we've

$$E_i = m v_{\perp 0i}^2, \quad E_f = \frac{5}{2} m v_{\perp 0i}^2$$

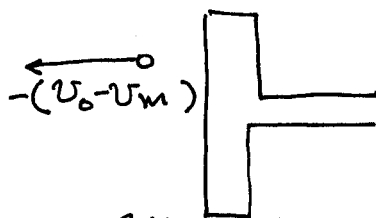
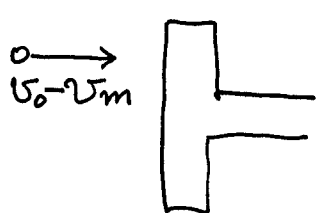
so, since

$$E_i = 1 \text{ keV}, \quad E_f = 2.5 \text{ keV}$$

b) → We've to first determine the change in velocity per bounce:



In the frame of the piston, we've



• Before collision

• After collision

(The collision is elastic and the mass of the piston \gg m_{proton})

so, in the lab frame,

$$v_f = v_f' + v_m = -(v_0 - v_m) + v_m = -v_0 + 2v_m$$

hence, the change in velocity is $2|v_m|$

→ Now let's figure out how many bounces we need for achieving $v_{\parallel 0f}$. since the change in momentum for each bounce is $\Delta \vec{p}_{\parallel} = 2m|v_m|$. So, for N bounces, we've

$$p_{\parallel f} = p_{\parallel i} + N \Delta p$$

$$\tau = \frac{\Delta x_{total}}{v_{||0}} = \frac{15.5L}{4.65 \times 10^5} = 3.33 \times 10^{-5} L$$

So $\Delta x_{total} = 15.5L$ &

do 15.5 bounces!

So you determine the time it takes to

$$= \frac{2}{3} v_{||} = 7.65 \times 10^5 \frac{m}{s}$$

$$v_{||0} = \frac{1}{2} (v_{||i} + v_{||f}) = \frac{1}{2} (v_{||i} + 2v_{||i})$$

Now, let's determine $v_{||i}$ (energy)

(so $L \approx$ constant)

the time it takes to reach $v_{||f}$ & escape.

Now, assume that $L \gg 25m$, where τ is

bounces

$$N \approx 15.5 \text{ times!}$$

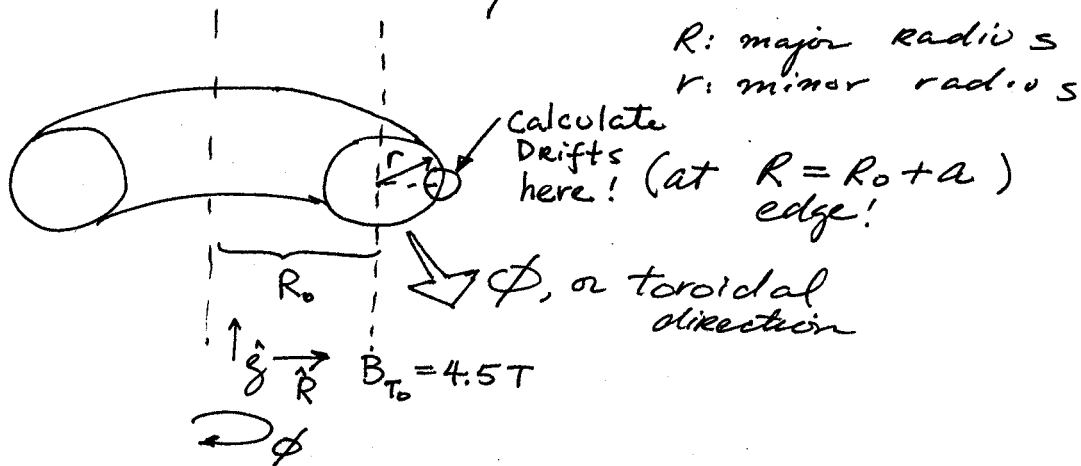
$$N = \frac{1}{2} \frac{3.1 \times 10^5 \text{ m/s}}{(10 \text{ km/s})} \left(1 \text{ keV} = \frac{1}{2} m_p (v_{||0}^2 + v_{||0}^2) \right) \text{ w/ } v_{||0} = 250!$$

$$\frac{\Delta p}{P_{||f} - P_{||i}} = N = \frac{v_{||f} - v_{||0}}{25m} = \frac{2v_{||0} - v_{||0}}{25m} = \frac{1}{2} \frac{2v_{||0}}{25m}$$

from pt. a

b) Const then,

4) Numbers for Drifts Tokamak Geometry



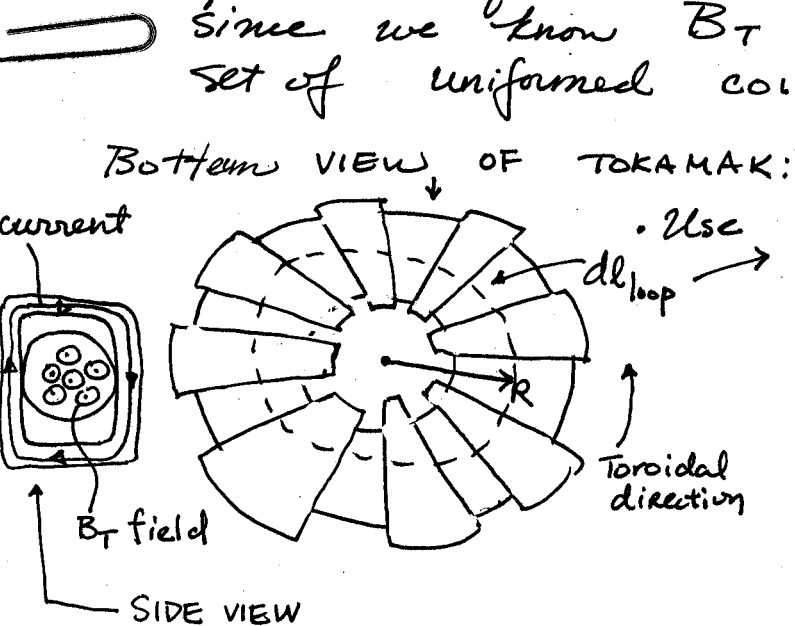
→ 1st thing we're to do is calculate B_T at the edge.

→ from class:

$$B(r) = \frac{B_0 R_0}{R} = \frac{4.5 T (.66 M)}{(.66 M + .2 M)} = 3.41 T$$

→ Derivation from Maxwell equation:

- We know $B \sim B_T = 4.5 T$ at $r=0$
- Using the low- β assumption, we can use Maxwell's equation to solve for $B_T(\hat{R})$, since we know B_T is created through a set of uniformed coils:



SIDE VIEW
 B_T field
 Toroidal direction
 R

4) Cont

$\vec{E} \times \vec{B}$ drifts:

$$\text{In SI, } \vec{v}_{\text{EXB}} = \frac{\vec{E} \times \vec{B}}{B^2} = \frac{E_r B_T (R+a)}{B_T (R+a)^2} = \frac{10^5 \text{ V/m}}{3.41 \text{ T}}$$

$$\boxed{\vec{v}_{\text{EXB}} = \frac{2.9 \times 10^4 \text{ m}}{\text{s}}}$$

Independent of q , m , or v_L !

$\vec{v} \perp \vec{B}$ drift

In SI,

$$v_{\text{VB}} = \pm \frac{1}{2} v_L r_L \frac{\vec{B} \times \nabla B}{B^2} \quad \text{where } r_L = \frac{m v_L}{q B}$$

Note:

$$\begin{pmatrix} \hat{r} \times \nabla B \\ \hat{r} \hat{\phi} \hat{z} \\ 0 B_T 0 \\ (v_B)_R 0 0 \end{pmatrix} = \hat{z} (-B_T \nabla B)_R = \frac{+B_T B_0 R_0}{R^2}$$

$$v_{\text{VB}} = \pm \frac{1}{2} \frac{v_L^2 m}{q B^3} \vec{B} \times \nabla B$$

$$v_{\text{VB}} = \pm \frac{1}{2} \frac{v_L^2 m}{q B_T^3} \left(\hat{z} \frac{B_T B_0 R_0}{R^2} \right)$$

$$v_{\text{VB}} = \pm \frac{1}{2} \frac{2 T B_0 R_0}{q B_T^2 R^2} \hat{z} \quad \left(\begin{array}{l} \text{since } \frac{1}{2} m v_L^2 \sim T \\ \frac{1}{2} m v_{\parallel}^2 \sim \frac{1}{2} T \end{array} \right)$$

total: $\frac{3}{2} T = v_{\text{th}}$

then, in dimensionless form,

$$\boxed{\left| \frac{v_{\text{VB} \perp \text{B}}}{v_{\text{th}i}} \right| = \frac{T B_0 R_0}{q B_T^2 R^2 v_{\text{th}i}} \hat{z} = \frac{T}{q B_T R v_{\text{th}i}} \hat{z}} \quad \text{for our problem}$$

$$v_{\text{th}i} = \left(\frac{3T}{m} \right)^{\frac{1}{2}} = 7.6 \times 10^5 \frac{\text{m}}{\text{s}} \quad \text{for hydrogen ions}$$

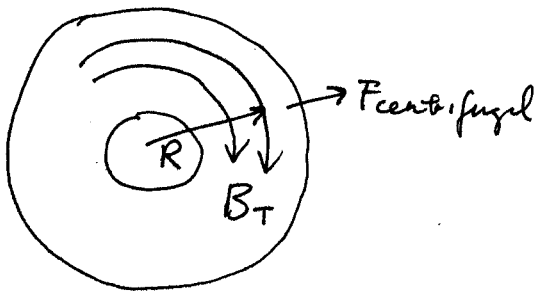
then

$$\left| \frac{v_{\text{DRLB}}}{v_{\text{thi}}} \right| = \frac{(2000 \text{ eV})(1.6 \times 10^{-19} \text{ J/eV})(4.5 \text{ T})(.66 \text{ m})}{(1.6 \times 10^{-19} \text{ C})(3.41 \text{ T})^2 (.66 \text{ m} + .21 \text{ m})^2 v_{\text{thi}}}$$

$$= 8.9 \times 10^{-4}$$

Curvature Drift

TOP VIEW



In SI Units

$$v_{\text{cur}} = \frac{m v_{\text{thi}}^2}{q B^2} \frac{\vec{R} \times \vec{B}}{R^2}$$

$$\vec{R} \times \vec{B} = \hat{z} (R B_T)$$

$$v_{\text{cur}} = \frac{T R B_T}{q B_T^2 R^2} \hat{z}$$

$$v_{\text{cur}} = \frac{T}{q B_T R} \hat{z}$$

for our geometry

In dimensionless form:

$$\frac{v_{\text{cur}}}{v_{\text{thi}}} = \frac{T}{q B_T R v_{\text{thi}}} \hat{z}$$

$$= \frac{(2000 \text{ eV})(1.6 \times 10^{-19} \text{ J/eV})}{(1.6 \times 10^{-19} \text{ C})(3.41)(.66 + .21 \text{ m})(7.6 \times 10^5 \text{ m/s})}$$

$$\frac{v_{\text{cur}}}{v_{\text{thi}}} = 8.9 \times 10^{-4}$$

→ Curvature & DRLB Drifts add!

→ only dependent on T & q