Introduction to Neural Computation

Prof. Michale Fee MIT BCS 9.40 — 2018

Lecture 12 - Spectral analysis II

Game plan for Lectures 11, 12, and 13 — Develop a powerful set of methods for understanding the temporal structure of signals

- Fourier series, Complex Fourier series, Fourier transform, Discrete Fourier transform (DFT), Power Spectrum
- Convolution Theorem
- Noise and Filtering
- Shannon-Nyquist Sampling Theorem
 - https://markusmeister.com/2018/03/20/death-of-the-sampling-theorem/
- Spectral Estimation
- Spectrograms
- Windowing, Tapers, and Time-Bandwidth Product
- Advanced Filtering Methods

• Some code

```
WSpec.m × recordaudio.m ×
                                continuous_cos.m 💥
                                                     +
 1
       2
                                % number of samples in time
2 -
       N=2048;
3
 4 -
                                % sampling interval
       dt=.001;
       Fs=1./dt;
                            % sampling frequency
5 -
       time=dt*[-N/2:N/2-1];
6 -
                                % timebase
7
       20
8 -
       freg=20.;
                                % frequency of sine wave in Hz
9 -
       y=cos(2*pi*freq*time);
10
11
       yshft=circshift(y,[0,N/2]);
                                        % First shift zero point from center to
12 -
13
                                             % first point in the array
       ffty=fft(yshft, N)/N;
                                         % Now compute the FFT
14 -
15
                                             % Now shift the spectrum to put zero frequency
16 -
       Y=circshift(ffty,[0,N/2]);
17
                                             % at the middle of the array
18
       %Compute the vector of frequencies
19
20 -
       df=Fs/N;
       Fvec=df*[-N/2:N/2-1];
21 -
22
       28
```

• Some examples – sine and cosine

$$y(t) = \cos(2\pi f_0 t)$$
 $f_0 = 20Hz$ Continuous_cos.m



• Some examples – sine and cosine

$$y(t) = \sin(2\pi f_0 t)$$
 $f_0 = 20Hz$ Continuous_sin.m



• Power spectrum of sine and cosine

Continuous_sin.m



• Some examples – square waves



Power spectrum – square wave



• Some examples – square waves



• Some examples – square waves

0 -1 E_ -0.8 -0.6 -0.4 0.2 0.6 0.8 -1 -0.2 0 0.4 Time (s) $f_0=10.75Hz$ 0.5 0 -0.5 -300 -200 -100 100 200 300 400 -400 0 Frequency (Hz)

Learning Objectives for Lecture 12

- Fourier Transform Pairs
- Convolution Theorem
- Gaussian Noise (Fourier Transform and Power Spectrum)
- Spectral Estimation
 - Filtering in the frequency domain
 - Wiener-Kinchine Theorem
- Shannon-Nyquist Theorem (and zero padding)
- Line noise removal

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Fourier transform pair

Square pulse $y(t) = \begin{cases} 1 \text{ if } |t| < \Delta T / 2 \\ 0 \text{ otherwise} \end{cases}$ Sinc function $Y(f) = \Delta T \, \frac{\sin(\pi \Delta T \, f)}{\pi \Delta T \, f}$ ΔT , $\Delta F \approx FWHM$ $\Delta F \approx \frac{1.2}{\Lambda T}$

Square_window.m



Fourier transform pair

The Fourier transform of a Gaussian

ls a Guassian!

$$\Delta F = \frac{1}{\Delta T}$$

Time-bandwidth product

 $\Delta T \Delta F = 1$



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Relation between Fourier transform and convolution





Convolution in the time domain

Multiplication in the frequency domain

$$y(t) = \int_{-\infty}^{\infty} d\tau \, g(\tau) x(t-\tau)$$

 $Y(\omega) = G(\omega)X(\omega)$

Fourier transform of a convolution

$$y(t) = \int_{-\infty}^{\infty} d\tau \, g(\tau) x(t-\tau)$$

$$Y(\omega) = G(\omega)X(\omega)$$

$$Y(\omega) = \int_{-\infty}^{\infty} dt \ y(t) \ e^{-i\omega t}$$

$$Y(\omega) = \int_{-\infty}^{\infty} dt \ \int_{-\infty}^{\infty} d\tau \ g(\tau) x(t-\tau) \ e^{-i\omega t}$$

$$= \int_{-\infty}^{\infty} d\tau \int_{-\infty}^{\infty} dt \ g(\tau) x(t-\tau) e^{-i\omega t}$$

$$= \int_{-\infty}^{\infty} d\tau g(\tau) \int_{-\infty}^{\infty} dt \ x(t-\tau) e^{-i\omega t}$$

$$= \int_{-\infty}^{\infty} d\tau g(\tau) \int_{-\infty}^{\infty} dt \, x(t-\tau) e^{-i\omega(t-\tau)} e^{-i\omega\tau}$$
$$= \int_{-\infty}^{\infty} d\tau g(\tau) e^{-i\omega\tau} \int_{-\infty}^{\infty} dt \, x(t-\tau) e^{-i\omega(t-\tau)}$$
$$= \int_{-\infty}^{\infty} d\tau g(\tau) e^{-i\omega\tau} X(\omega)$$
$$= G(\omega) X(\omega)$$

Relation between Fourier transform and convolution

Fourier Transform Pairs

 $y(t) \Leftrightarrow Y(\omega)$ $g(\tau) \Leftrightarrow G(\omega)$ $x(t) \Leftrightarrow X(\omega)$

Convolution in the time domain

Multiplication in the frequency domain

$$y(t) = \int_{-\infty}^{\infty} d\tau \, g(\tau) x(t-\tau)$$

$$Y(\omega) = G(\omega)X(\omega)$$

Convolution in the frequency domain

Multiplication in the time domain

$$Y(\omega) = \int_{-\infty}^{\infty} d\omega' G(\omega') X(\omega - \omega')$$

$$y(t) = g(t) x(t)$$

Using the Convolution Theorem

Gaussian-windowed cosine

Cos_Gauss_pulse.m



Using the Convolution Theorem Square-windowed cosine g(t) = square $x(t) = cos(2\pi f_0 t)$ cos_Gauss_pulse.m



$$y(t) = g(t)x(t)$$

Product in the time-domain

$$Y(f) = G(f) * X(f)$$

Convolution in the frequencydomain!

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Gaussian noise

 Each sample drawn independently from a Gaussian distribution y=randn(1,N); white_noise.m



Fourier transform of Gaussian noise

• Some examples – Gaussian white noise white_noise.m



The Fourier transform of Gaussian noise, is just Gaussian noise!

Power spectrum of noise

• Is very...



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- Say we want to find the spectrum S(f) of a signal y(t).
- Often we only have short measurements of y(t) (e.g. trials)



We can just average!

$$\hat{S}(f) = \frac{1}{N} \sum_{i=1}^{N} \hat{S}_{i}(f)$$

- The same principle applies to longer signals.
 - Break the signal into shorter pieces
 - Compute the power spectrum in each window



- We could just take the FFT of each piece.
 - But we know that a 'square windowing' means that the spectrum becomes convolved with the spectrum of the square window!

• We will multiply each window by a smooth function called a 'taper'.



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- A common problem is to find a small signal in noise
 - This can be a challenge



Power spectrum of noise

But the average spectrum of white noise is flat!



The power spectrum is a <u>power spectral density</u>. It tells us the amount of variance per unit frequency. Variance is power density times bandwidth.

$$\sigma^2 = 0.002 \frac{V^2}{Hz} x \ 500 Hz = 1$$
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Low-pass filtering



Convolution as a filter

• If the convolution is equivalent to a multiplication in the frequency domain...

$$Y(\boldsymbol{\omega}) = G(\boldsymbol{\omega})X(\boldsymbol{\omega})$$

what does this do to the power spectrum?

$$S(\omega) = |Y(\omega)|^2 = |G(\omega)|^2 |X(\omega)|^2$$

• The power spectrum of the filtered signal is just the power spectrum of the original signal times the spectrum of the filter!

Power spectrum of filtered noise

Power spectrum of white noise convolved with a Gaussian window

x(t)



Y(f) = G(f)X(f)

 $|Y(f)|^{2} = |G(f)|^{2} |X(f)|^{2}$





Autocorrelation of filtered noise

Autocorrelation



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Wiener-Kinchine theorem

• The power spectrum and autocorrelation functions are related by Fourier Transform

Autocorrelation



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• Remember that frequency is discretized...



• Which means that our function is periodic in time!



• But time is also discretized...



• Which means that our FFT is periodic in frequency!



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... and the separation between the spectra in the frequency-domain is given by the sampling rate.

• Sampling rate must be greater than twice the bandwidth of the signal $F_{samp} > 2B$.



Sampling rate too low



- If the sampling rate is greater than twice the bandwidth of the signal $F_{samp} > 2B$
- Then you can perfectly reconstruct the original signal.

 Not just at the sampled points, but continuously, at every point!



- Remember the Convolution Theorem...
- Multiply the periodic Fourier transform by a square window in the frequency domain...



• This convolves the time-domain signal with a Kernel that is the Fourier transform of the square window... In other words, we can smooth it, so that it is no longer sampled. The smoothed signal has the same spectrum as the sampled signal!

Zero-padding

• If $F_{samp} > 2B$ you can interpolate the values of the function y(t) with arbitrarily high temporal resolution by zero padding.





Y=fft(y, N)/N; zero_pad_factor=4.; Nresamp=N*zero_pad_factor; Yresamp=zeros(1,Nresamp); Yresamp(1:N/2)=Y(1:N/2); yRes=2*real(ifft(Yresamp)*Nresamp);

% Compute the FFT

% new number of frequency bins
% fill a new array with zeros
% insert the positive frequency part
% compute the inverse FFT

Zero-padding

Or Fun with FFT and IFFT

- Zero padding in the time domain gives finer spacing in the frequency domain.
- Just add zeros to the end of a tapered window before you FFT. The spectrum that you get is the same, but it has samples that are more finely spaced. (No higher frequency resolution though.)
- MATLAB[®] makes this easy...

```
zero_pad_factor=4.;
y=Data;
Nfft=N*zero_pad_factor;
Y=fft(y, Nfft);
```

% Data is a vector in time with N samples
% number of points you want in the spectrum
% Compute the FFT
% The array Y will now have 4 times the samples

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Line noise removal

• Another common problem is to remove a small periodic noise in your signal.



Frequency Bin

- While the periodogram is a terrible spectral estimator for nonperiodic broadband signals, it is a great estimator for perfectly stationary single-frequencies... like contamination from 60Hz.
- So, if you have a single offending frequency component...

Off with its head!

Line noise removal

• Just find those lines in Y(f) and set them to zero!



• Then inverse FFT Y(f) to get the cleaned up signal...



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